

The evidence of latitudinal asymmetry of the ammonia absorption on Saturn

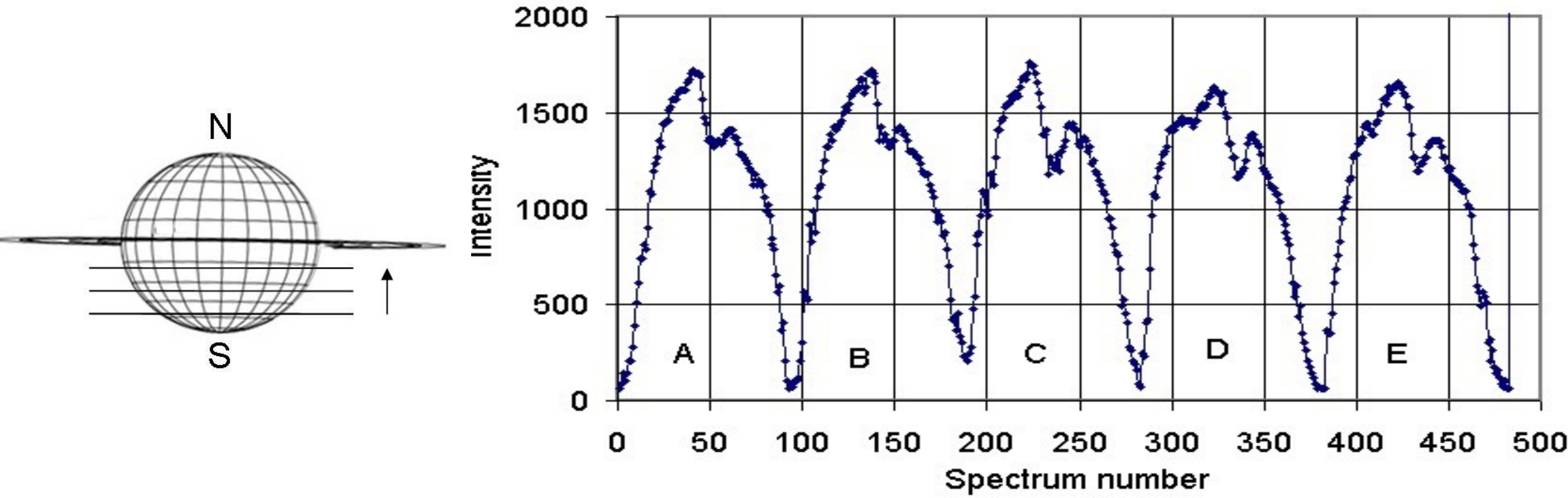
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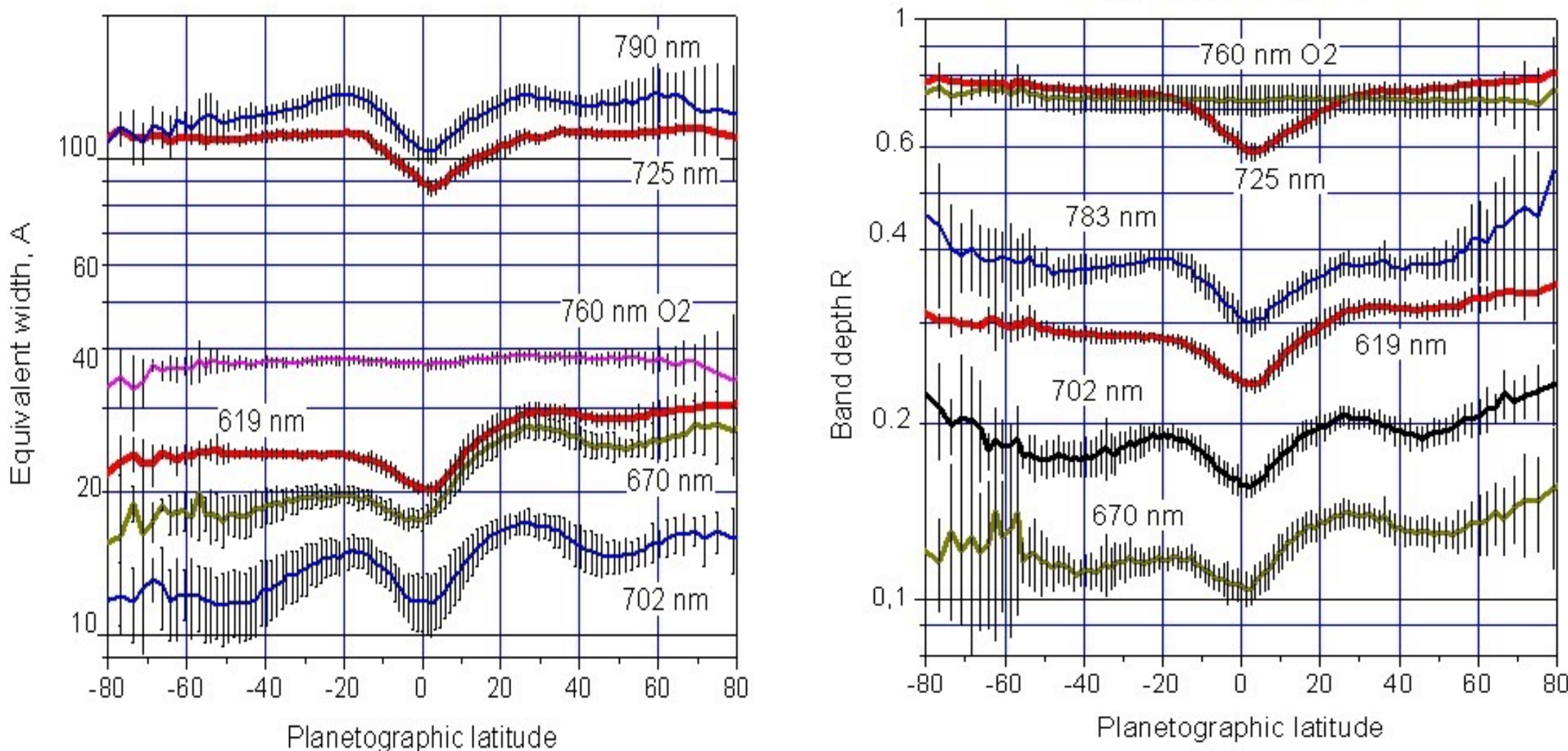
Saturn in the equinox-2009

Our research had been directed to the study of the methane and ammonia. Absorption bands behaviour on different latitudes and comparison of their intensity at Southern and Northern hemispheres. This period is transitional between primary insolation of S-hemisphere in preceding years and the growth of insolation of N-hemisphere during next seven years. We continue the spectrophotometric scanning of Saturn's disk.



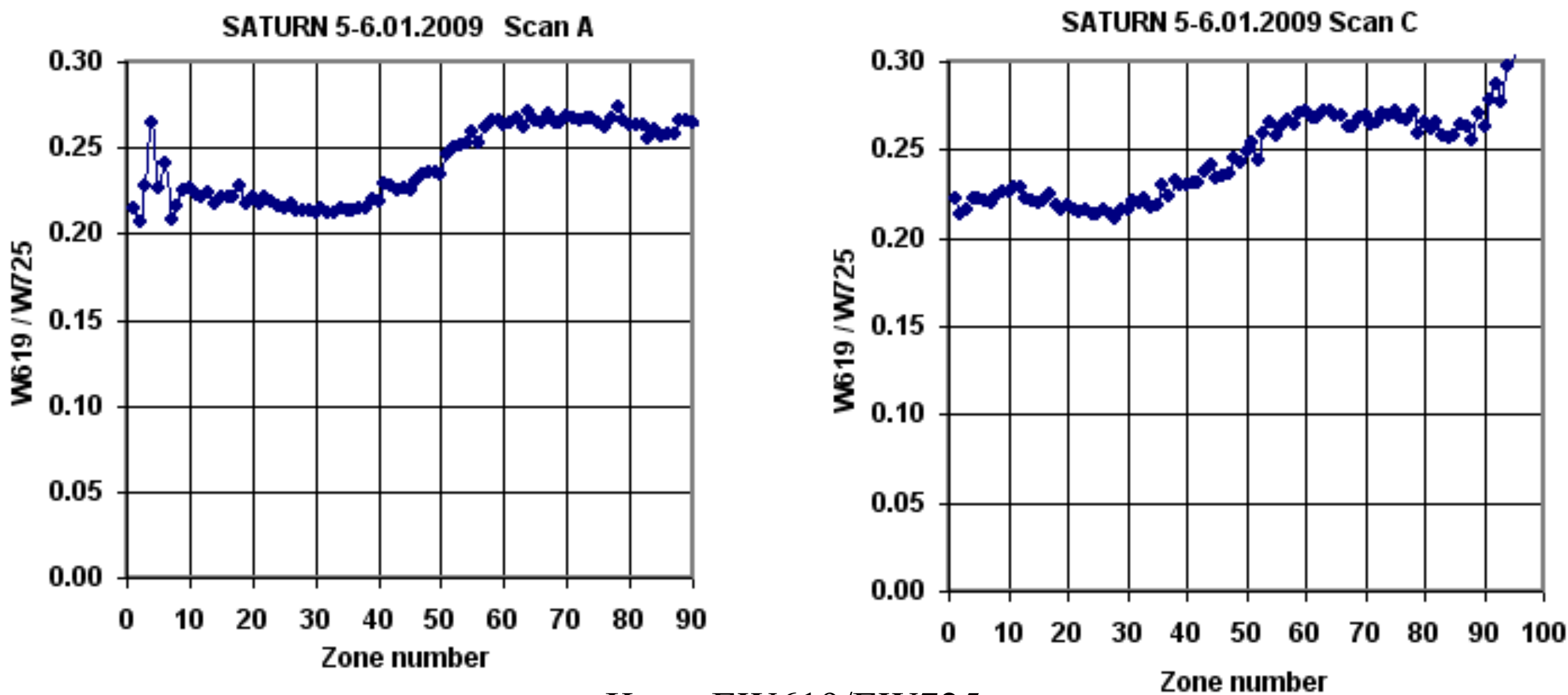
South - North scanning of Saturn for recording zonal spectra (90-100 spectra per scan)

Latitudinal variations of the methane absorption



There are the latitudinal variations of the CH₄ absorption bands equivalent widths and central depths in logarithmic scale. They were derived from the averaging of 5 scans data in 5 January 2009.

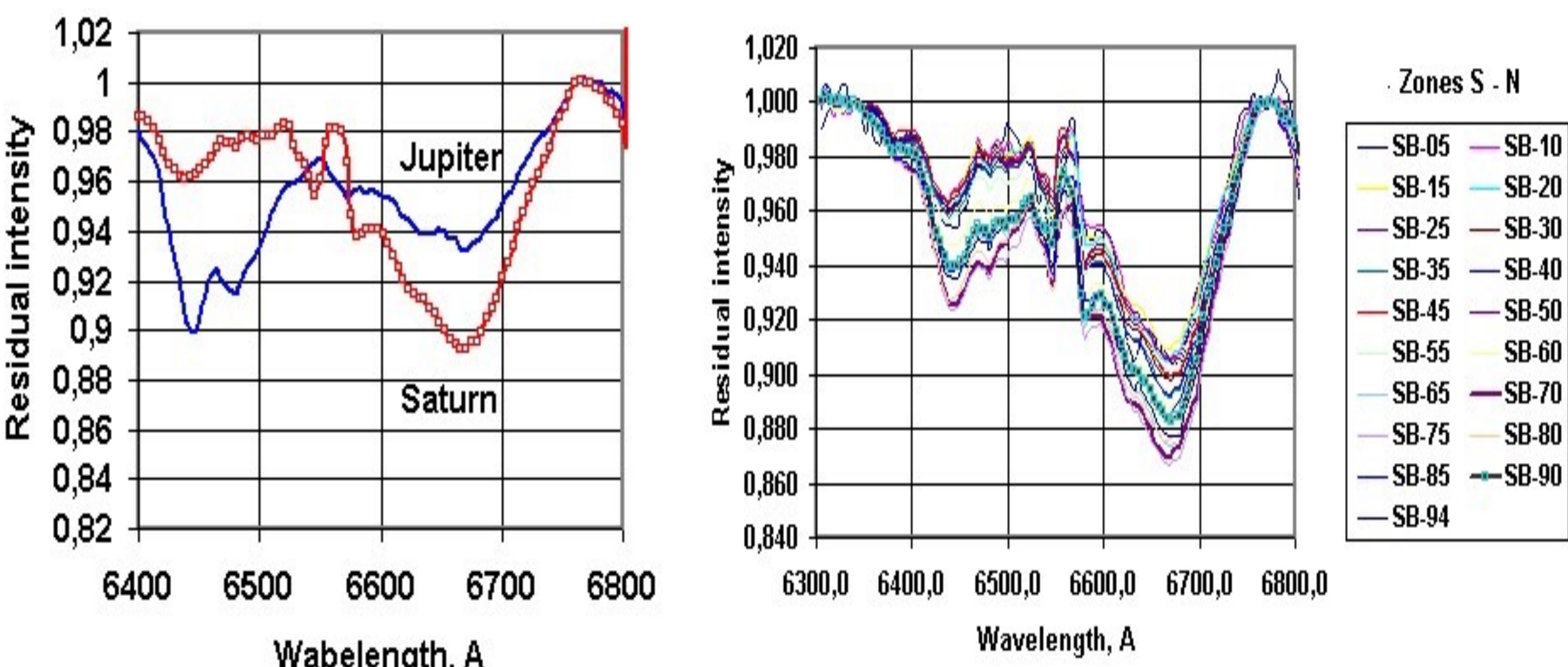
The south - north differences of the relative absorption in 619nm and 725nm methane bands



K_w = EW₆₁₉/EW₇₂₅

In 2008-2009 K_{ws}=0.215 and K_{wn}=0.265.
In 2010 an amplitude of S-N differences noticeably decreased: K_{ws}=0.225 but K_{wn}=0.250.
In the beginning of 2008 the temperate latitudes showed extreme K_w values but in the next seasons the latitudinal variations of K_w within one hemisphere were smoothed.

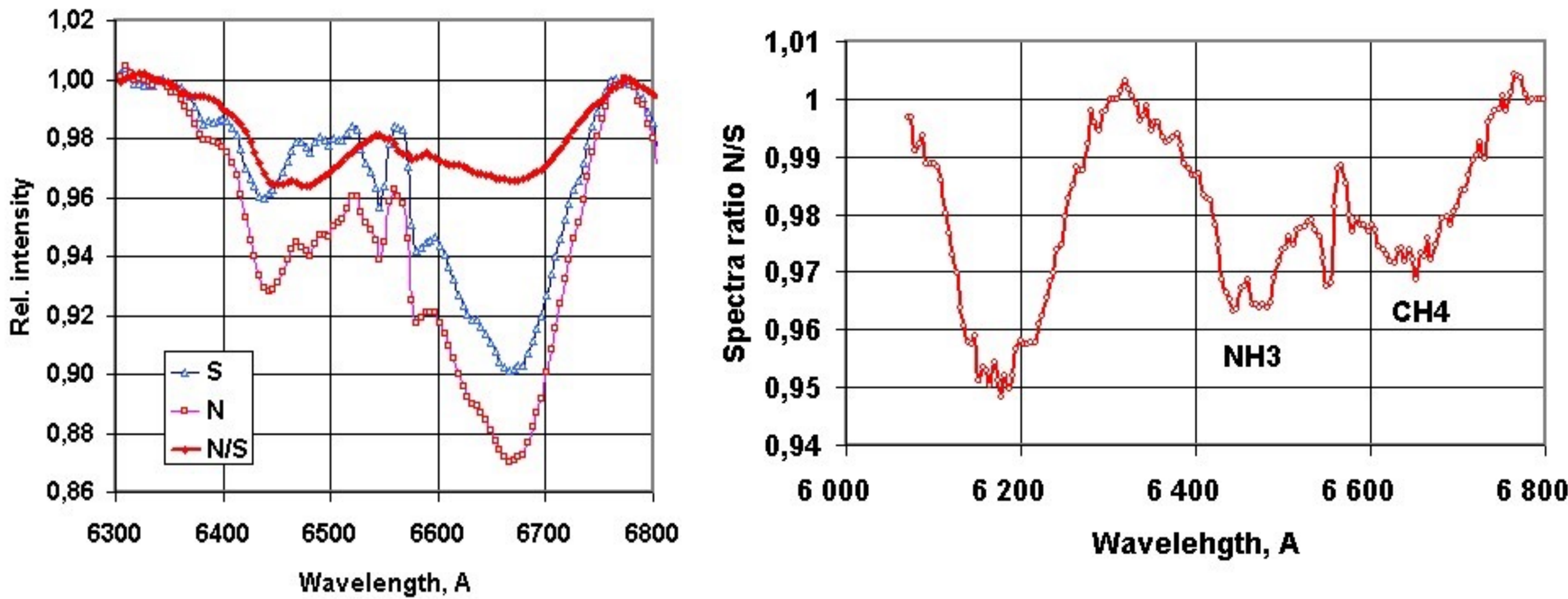
The ammonia and methane absorption band NH₃ 645nm + CH₄ 678nm



Jupiter and Saturn

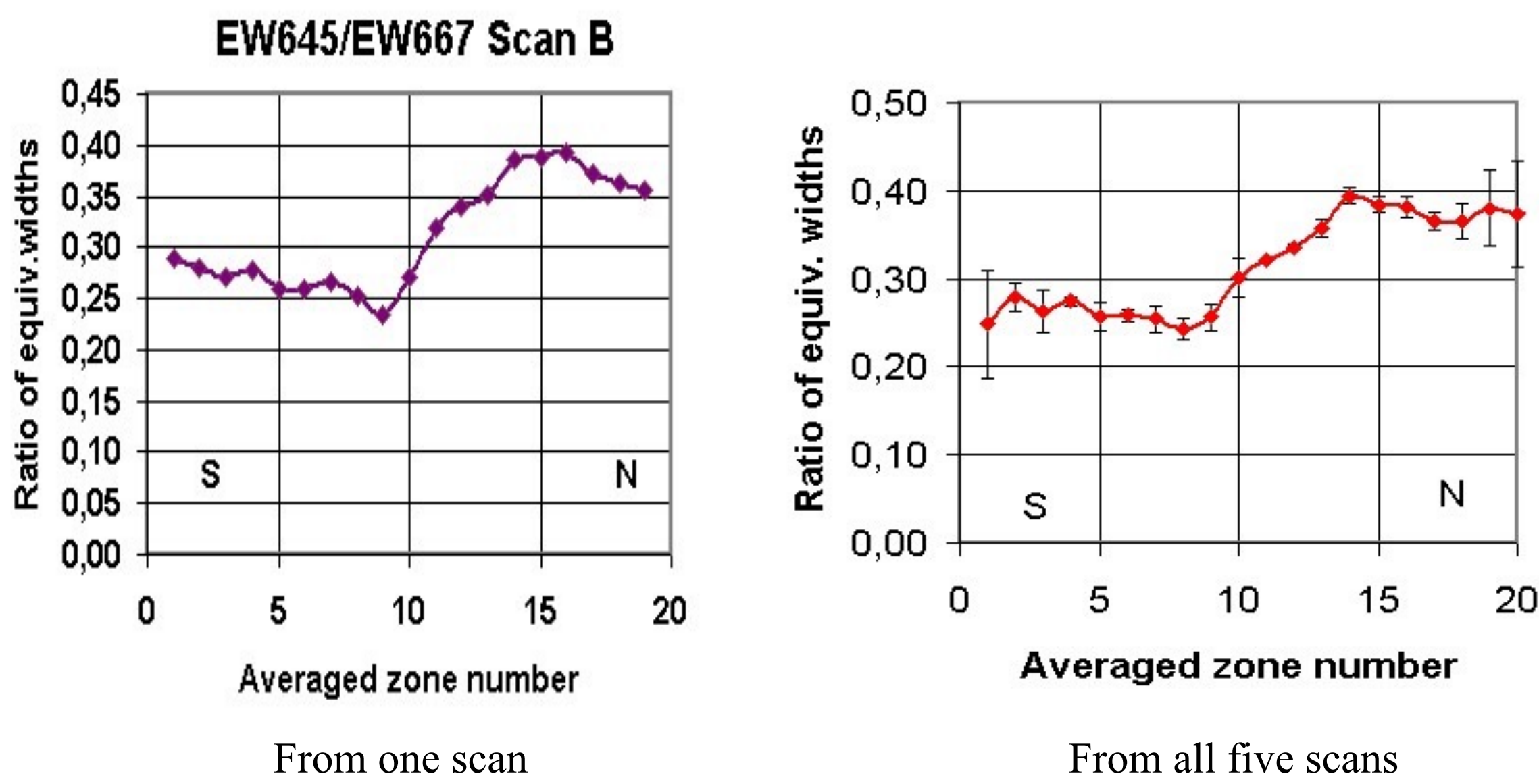
Saturn – zonal spectra

A comparison of the NH₃ 645nm + CH₄ 678nm absorption band profiles for S- and N- temperate latitudes and their ratio



The ratio South/North for CH₄ 619 nm and NH₃ 645+CH₄ 678 nm bands profiles

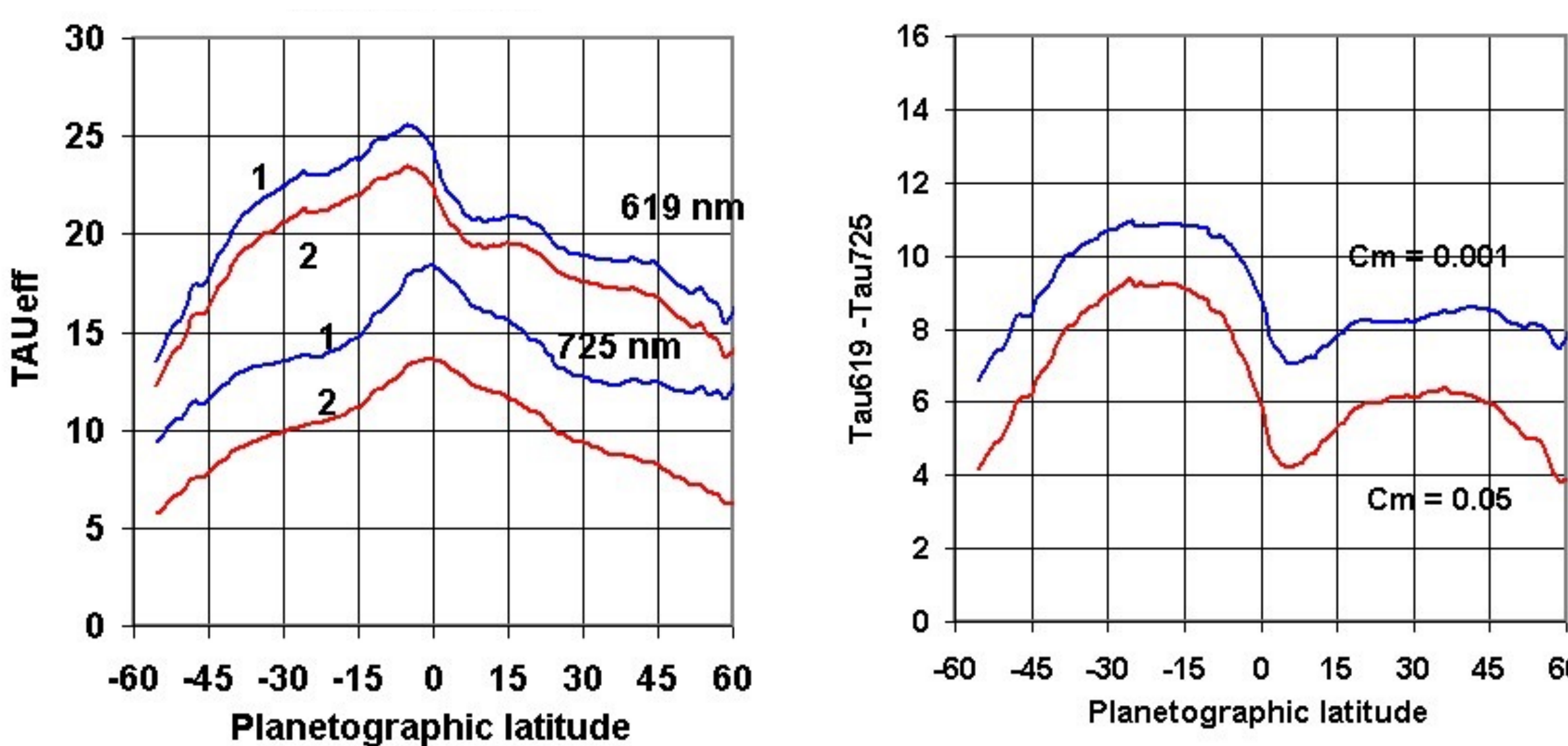
The profiles and equivalent widths ratios of the NH₃ 645nm and CH₄ 667nm absorption bands



From one scan

From all five scans

Preliminary estimates of the absorption bands formation effective optical depths for CH₄ 619nm and 725nm and their differences at two methane abundance C_m [m/amagat]



The results show that the increased NH₃ absorption in Saturn's northern hemisphere coincides with the increase of relatively weak CH₄ absorption bands, observed also in the northern hemisphere. At the same time stronger absorption band of methane, for example, 725 nm band, the similar hemispheric difference is not detected. This may be due to a decrease of the volume density and the aerosol scattering coefficient on the large effective optical depths by raising the temperature. In the upper part of the cloud cover the difference in density of the cloud layer apparently absent or much less pronounced.